Finding Failed Supernovae

Scott Adams OSU/Caltech July 26, 2016



Evidence for FSNe

The Mapping Between SN Progenitors and Outcomes

- ~ 10 with reasonable multi-band photometry, almost all for Type IIP
- There is a deficit of higher mass progenitors (Kochanek et al. 2008)
- Best quantified by Smartt et al. (2009) Type IIP red supergiant progenitors seen from ~8 to ~ $17M_{\odot}$, but expected to die as a red supergiants up to ~ $25M_{\odot}$
 - See Smartt 2015 for recent review





- Distribution of progenitors essentially cuts off exactly where the compactness jumps upwards
- Missing red supergiant progenitors would correspond to 10-30% of core collapses







Other Evidence

Same high mass progenitor deficit when estimating masses from nearby stellar population (Jennings et al. 2014)

Compact object mass function (Kochanek 2014, 2015)

SN rate may not match massive star formation rate (Horiuchi et al. 2011, 2014)

Massive merging black holes detected by LIGO (Abbott et al. 2016)





Associated Transients?



Little studied:

- Red supergiants → ejected envelope powers a ~ 10⁶L_☉, cool (~3000-4000K), long (~ year) transient powered largely by H recombination (Lovegrove & Woosley 2013), along with a ~10 day, more luminous, shock breakout peak (Piro 2013), and very late time dust formation (Kochanek 2014).
- Can have direct collapse with no transient (Woosley & Heger 2012)
- Accretion may power a short, luminous transient (Kashiyama & Quataert 2015)

Focus on disappearance rather than relying on transient signal





The Nadezhin (1980) Mechanism

- The envelopes of red supergiants are so weakly bound that the drop in gravitational potential during a failed SN through neutrino mass-energy loss unbinds the hydrogen envelope
- Numerically verified by Lovegrove & Woosley (2013)

→ Failed SN of red supergiants (and WR stars) produce black holes with the mass of the helium core not the total mass of the star

→ Failed ccSN provide the first natural explanation of the compact remnant mass function – successful SN (almost always) make NS with negligible fall back, failed SN make BH with the mass of the helium core (Kochanek 2014)



Nadezhin Mechanism









How Do You Find Failed SN?

Cleanest would be gravitational waves or neutrinos – but only possible in Galaxy, so rate is ~1/500 years

Instead search nearby galaxies for stars which "disappear", possibly with an intervening transient (Kochanek et al. 2008)

Feasible on an 8m to ~ 10 Mpc – galaxies with an SN rate of ~1/year \rightarrow failed SN every 3-10 years



LBT Survey for Failed Supernovae Kochanek et al. 2008; Gerke et al. 2015; Adams et al. in prep.



LBT Survey



- 8.0 27 nearby galaxies
- •6.0 Historical ccSN rate of ~ 1/year → failed SN every 3-10 years
 - UBVR imaging to a typical depth of about 1 count per $\rm L_{\odot}$
 - Analyze using difference imaging
 - Examine everything that:
 - varied by more than $10^4 {\rm L}_{\odot}$ in any band
 - > $10^5 L_{\odot}$ for 3 months to 3 yrs





FSN Candidate: N6946-BH1



N6946BH1: light curve

Global Relay of Observatories Watching Transients Happen





Global Relay of Observatories Watching Transients Happen HST Follow-up





N6946-BH1: Dusty Wind





N6946-BH1: Dusty Wind



N6946-BH1: Dusty Shell





N6946-BH1: Dusty Shell





Bolometric Luminosity Evolution







Bolometric Luminosity Evolution







Bolometric Luminosity Evolution







Failed SN Fraction







Failed SN Fraction



























- N6946-BH1
 - Possibly first failed SN & first newly-formed BH discovered
- If confirmed: ~14% of core collapses fail (5-47%)
 - Resolves RSG problem; explains BH mass function
 - ZTF should detect dozens fSN SBO per year
 - Timescales & peak wavelength ideal for GROWTH follow-up
 - SBO constrains progenitor radius
 - SBO + recombination plateau together constrain explosion energy and ejecta mass
- If rejected: <35% (90% confidence) of core collapses fail





N6946BH1: light curve



